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GPU Parallelization of Spin-Tracking Simulations for the nEDM@SNS Experiment

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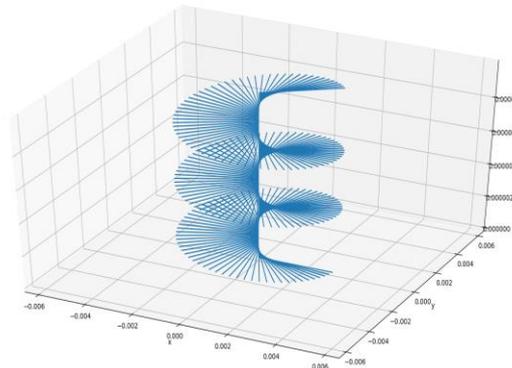
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Overview

- The nEDM@SNS experiment extracts the EDM from the precession of ^3He atoms and spin-dependent signal of neutrons capturing on ^3He
- Spin-tracking simulations are very computationally expensive
 - We need these simulations to better understand systematic effects that can lead to a false EDM measurement
 - Need fast spin-tracking simulations for 10^{11} events to have comparable sensitivity to expected experiment precision
 - Use Julia and CUDA to run simulations on GPU to utilize parallelization

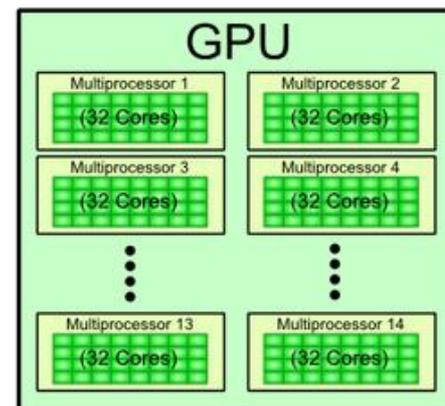
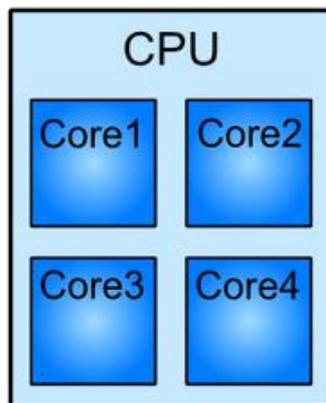


Spin precession of neutron with gravity



Simulations on GPUs

CPU	GPU
Central Processing Unit	Graphics Processing Unit
1-64 cores	100 - 7000 cores
Linear Processing	Parallel Processing
A handful of operations very rapidly	Thousands of operations at once



Spin Precession

- Spin precession described by the Bloch equation

$$- \dot{\sigma} = \gamma_n \sigma \times B - \frac{2d_n}{\hbar} \sigma \times E$$

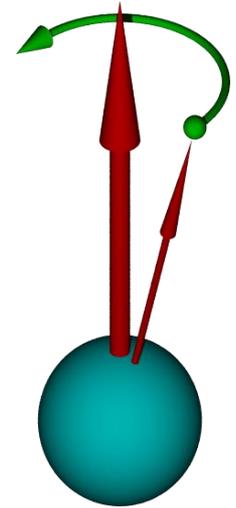
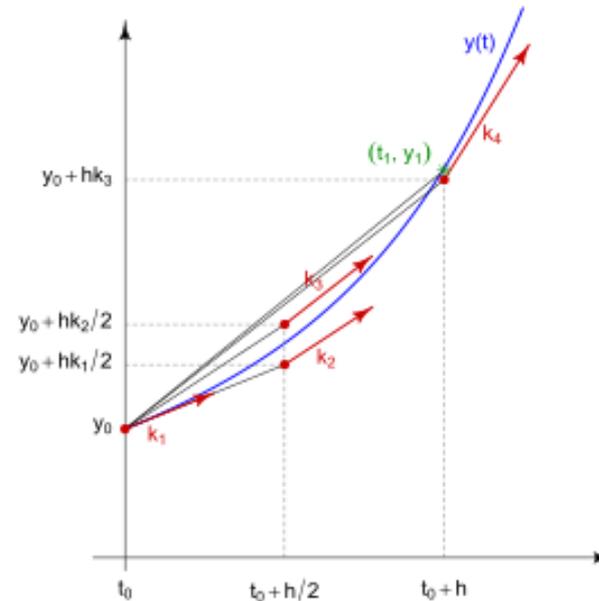
- Numerically integrated using 5th order Runge-Kutta method

- Used Cash-Karp parameters

$$- \frac{dy}{dt} = f(t, y)$$

$$- k_i = h f(x_n + a_i h, y_n + \sum_{j=1}^{i-1} b_{ij} k_j)$$

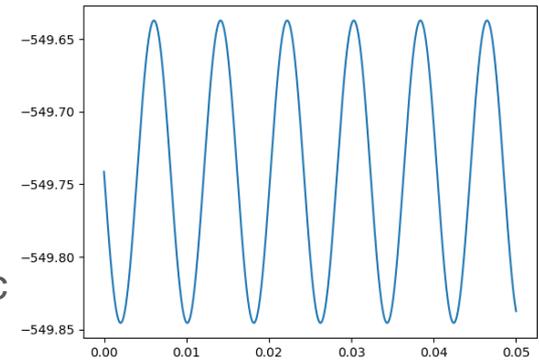
$$- y_{n+1} = y_n + \sum_{i=1}^6 c_i k_i$$





Other Integration Methods

- Symplectic integration
 - More stable than Runge-Kutta over long durations
 - Used for Hamiltonian where $\mathcal{H} = T(p) + V(q)$
 - Great for UCN in gravitational/magnetic fields
 - Reflections/variable step size become problematic
 - $\frac{dx}{dt} = f(t, v)$ and $\frac{dv}{dt} = g(t, x)$ satisfy the Hamiltonian
 - $\mathcal{H} = -\mu_n \sigma \cdot B - d_n \sigma \cdot E$
 - We investigated the transformation $\sigma = x$ and $\dot{\sigma} = v$
- Turn RK/SI step into a matrix equation
 - The differential equation must be linear (like the Bloch equation)
 - Matrix operations are fast on GPUs
 - Adding physics (reflections, etc.) slows it down
 - Time dependence is difficult
 - Magnetic field nonuniformities cause problems

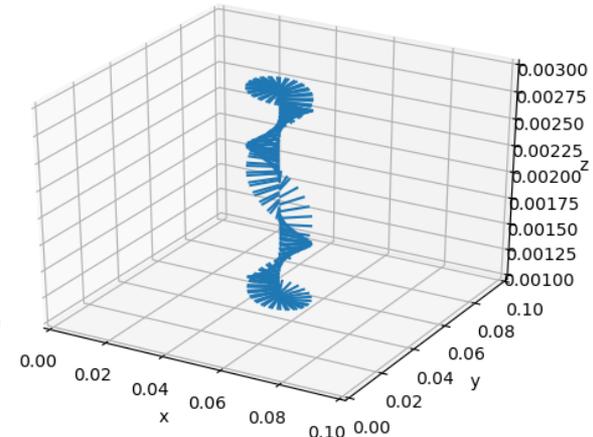


Spin precession frequency vs time with symplectic integration

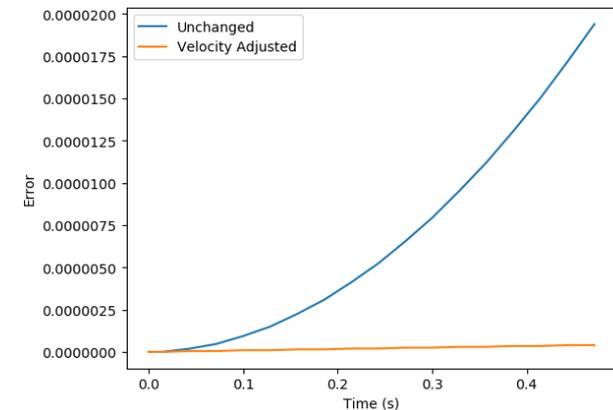


Wall Interactions

- Wall reflections are difficult
 - Magically change neutron's direction in the middle of a solver step
 - To avoid errors, you have to change the step size to reach the wall exactly
- Side walls are periodic to avoid variable step size
 - Achieved with a GPU kernel
- Problem: top/bottom walls break symmetry because of gravity
- Solution: duplicate measurement cell
 - Top cell has normal gravity
 - Bottom cell has "antigravity"
- Still run into problem of wrong gravity during part of step
 - Solved by scaling velocity



Neutron oscillating between top and bottom cells



Error accumulation with and without adjusting velocity after crossing top/bottom wall



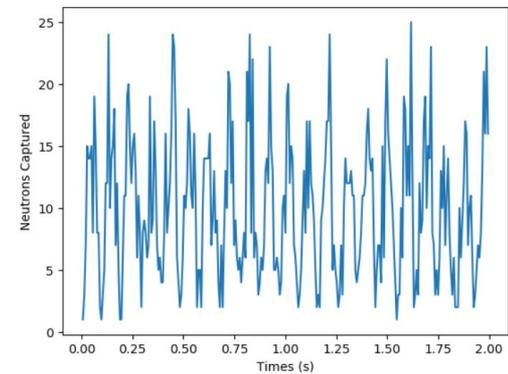
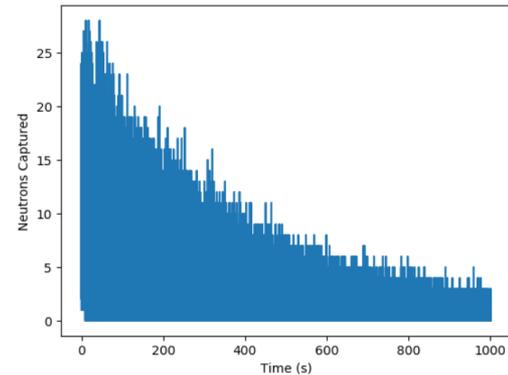
Wall Interactions

- Probability of diffuse reflections $\propto \cos \theta_i$
 - $\approx 20\%$ for normal incident angle
- Diffuse reflections must preserve detailed balance
 - For random number $y \in [0,1]$, $\theta = \cos^{-1}[(1 - y)^{1/3}]$
 - Azimuthal angle ϕ can be chosen from uniform distribution in 2π
- Wall losses added with constant probability to match expected lifetime of 2000s
 - Lifetime will be experimentally determined for each measurement cell



$n+^3\text{He}$ capture and β decay

- At each step, probability of $n+^3\text{He}$ capture event is dependent on relative spin between UCN and ^3He
 - $h\Gamma_3(1 - \cos \theta)$
- Probability of β decay is constant
 - $h\Gamma_\beta$
- Γ_3 and Γ_β come from lifetimes
 - $\Gamma_3 = \frac{1}{\tau_3} = \frac{1}{500 \text{ s}}$ $\Gamma_\beta = \frac{1}{\tau_\beta} = \frac{1}{881.5 \text{ s}}$
- Plot binned events vs time
- Large-scale decay caused by decrease in number of neutrons in the cell
- Zooming in shows oscillation
 - Frequency of oscillation with ^3He precession signal used to determine nEDM

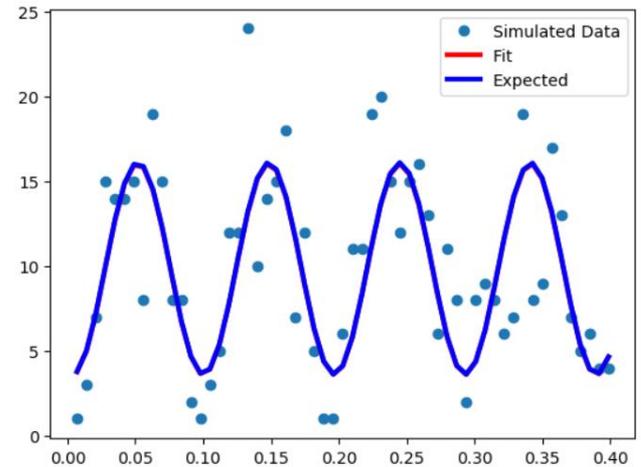


Simulated scintillation signal vs time



Simulated Signal Fit & Analysis

- Rate of events is
 - $\Gamma(t) = N(t) \times [\Gamma_\beta + \Gamma_3(1 - \cos[\omega t + \phi_0])] + \Gamma_B(t)$
 - $N(t) = N_0 e^{-t/\tau_{eff}}$
- Maximize log likelihood to obtain fit parameters
- $\chi^2 \approx -2 \log \Lambda$
 - $\Lambda = \mathcal{L}/\hat{\mathcal{L}}$ is the maximum likelihood ratio
- $\log \Lambda = \sum_{i=1}^{N_{bins}} [k_i \log\left(\frac{\lambda_i}{k_i}\right) - (\lambda_i - k_i)]$
- $\omega = (\gamma_3 - \gamma_n) \frac{\omega_3}{\gamma_3} + \frac{2d_n}{\hbar} E$
- $1\sigma \approx 2.5 \times 10^{-23} e \cdot cm$

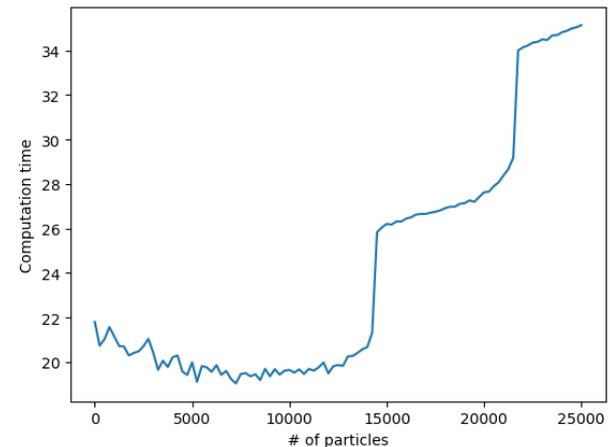


Maximum likelihood fit of simulated scintillation signal



GPU Computation Time

- Simulate $\approx 15,000$ particles simultaneously per GPU
- At $\approx 15,000$, another set of streaming multiprocessors is required for the task, causing a jump in computation time
 - Will remain approximately constant until another set is required
- Preliminary computation time per particle per iteration
 - CPU: $2 \mu\text{s}/\text{CPU}$
 - GPU: $8 \text{ ns}/\text{GPU}$ (with 225,000 particles)



Computation time vs number of particles simulated



Summary and Future Work

- GPUs look promising for studying computationally expensive systematic effects
- Investigating additional physics
 - Realistic B-field maps
 - Time-dependent systematics
 - Tracking of ^3He particles
- Optimize GPU code to get better performance
- Compare with CUDA C
- Goal: run hybrid CPU/GPU version on Summit supercomputer





Thank You!

Questions?

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